

COMPARISON OF AEROSOL OPTICAL DEPTH FROM FIVE SOLAR RADIOMETERS
DURING THE FALL 1997 ARM INTENSIVE OBSERVATION PERIOD

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Comparison of Aerosol Optical Depth from Five Solar Radiometers During the Fall 1997 ARM

Intensive Observation Period

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ABSTRACT

In the Fall of 1997 the Atmospheric Radiation Measurement (ARM) program conducted an Intensive Observation Period (IOP) to study aerosols. Five sun-tracking radiometers were present to measure the total column aerosol optical depth. The World Meteorological Organization in a 1993 workshop recommended a comparison of tracking sun radiometers on a high mountain site. This *de facto* comparison performed on the Southern Great Plains (SGP) rather than in a mountain setting nevertheless demonstrates the capabilities and limitations of modern tracking sunphotometers at a location typical of where aerosol measurements are required. The key result was agreement in aerosol optical depth measured by 4 of the 5 instruments within 0.015 (rms). The key to this level of agreement was meticulous care in the calibrations of the instruments.

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1. INTRODUCTION

In the last World Meteorological Organization workshop held to discuss the measurement of atmospheric optical depth (WMO 1993) it was recommended that the existing program to measure aerosol optical depth (AOD) sponsored by the WMO, that is, the Background Air Pollution Monitoring Network (BAPMoN) program, be discontinued and that the data that it had collected be regarded as suspect. The principal problem was that the instruments were unstable with regard to calibration, and the measurement protocol was such that too few measurements were made to try to discern the changes in calibration through use of Langley analyses.

New aerosol measurements under the auspices of the WMO were to be part of the Global Atmospheric Watch (GAW) program. Two of several specific recommendations for this new program in aerosol optical depth measurements were that automated sun-tracking would be required and that instruments to be used in the GAW program should be compared at a high altitude site.

In the Fall of 1997 the ARM program organized an IOP to study aerosol optical and physical properties using both remote sensing and in situ techniques. Five solar tracking radiometers for measuring total optical depth were present for the IOP leading to this *de facto* comparison along the guidelines of that recommended by the WMO. Although the comparison was not conducted at a high mountain site, there were several clear and stable days during the IOP.

The measurements were made between 15 September and 5 October 1997 at the SGP ARM central facility near Lamont, Oklahoma (36° 36' N, 97° 22' W, 316 m above sea level). Very clean to moderate aerosol loading was experienced over the three-week period. Mid-visible AOD values ranged from 0.025 to 0.3.

2. INSTRUMENTATION

The NASA Ames Research Center deployed their six-channel Ames Airborne Tracking Sunphotometer (AATS-6) at the SGP central facility of ARM for this IOP. This instrument, described by Matsumoto et al. (1987), uses an active sun sensor to keep the instrument pointed at the solar disk. The

central wavelengths and full widths at half maximum (FWHM) for the filters, are given in Table 1. The Si detectors are held at a constant temperature of 45 ± 0.6 °C. The field-of-view (FOV) of AATS-6 is 4.5° . A measurement sequence was repeated every 12 seconds with all filters scanned nine times then averaged in the first three seconds of the 12-second period. The instrument was calibrated by averaging the results of 6 successful morning Langley plots (Schmid et al. 1998) performed at the Mauna Loa Observatory in Hawaii ($19^\circ 32'$ N, $155^\circ 34'$ W, 3397 m above sea level) about two weeks before the IOP. A Langley plot performed during the IOP at the SGP site on the morning of September 29 yielded calibration constants that agreed within 0.5% with the Mauna Loa results.

Pennsylvania State University (Penn State) deployed an early generation sunphotometer fabricated at the University of Arizona (Shaw et al. 1973). After the IOP it was discovered that the filter wheel of that instrument had slipped slightly in the drive shaft and therefore the filters would not necessarily be perfectly aligned with the optical axis. Indeed the AOD obtained from that instrument revealed a positive bias and considerable scatter when compared to the other 4 instruments. Therefore we have excluded the data of the PSU sunphotometer from the comparisons shown below.

At the ARM SGP central facility a CIMEL sunphotometer/radiometer measures AOD. This instrument is also part of AERONET, a worldwide network of CIMEL sunphotometers (Holben et al. 1998). The CIMEL CE-318 points to the sun based on an ephemeris calculation and then fine tunes the pointing with an active sun sensor adjustment. Samples consist of triplets of measurements with each member of the triplet beginning 30 seconds apart and consisting of eight filter measurements completed within eight seconds; the triplets are repeated at every quarter air mass between two to five air masses and every 15 minutes when the air mass is less than two. The central wavelength and FWHM for each filter is given in Table 1. The field-of-view is 1.2° . The temperature of the instrument is monitored but not controlled. Calibration is based on a transfer of the calibration from an instrument that has recently been calibrated using the Langley technique at the Mauna Loa Observatory.

The multi-filter rotating shadowband radiometer (Harrison et al. 1994; MFRSR) has a hemispherical field-of-view. A band is positioned to alternately move completely out of the field-of-view

and then to block the sun according to a solar hour angle calculation allowing a measurement of the total downward and diffuse downward irradiance. The difference between the two measurements is the direct solar component normal to the receiver, and the direct normal component is calculated by dividing by the cosine of the solar-zenith angle and correcting for the angular response of the quasi-Lambertian detector. Sampling is every 20 seconds. The central wavelength and FWHM for each filter is given in Table 1. The temperature is held at 40°C. Calibration was based on a robust estimate using the 20 nearest successful Langley plots at SGP.

The rotating shadowband spectroradiometer (Harrison et al. 1999; RSS) has a Lambertian receiver and a shadowing sequence similar to the MFRSR, however, the detector is a 512-element photodiode array that receives its light input from the focus of a prism spectrograph. Sampling is performed once each minute. The spectral resolution between 350 and 1050 nm varies between 0.3 to 8 nm because of the prism dispersive element. The temperature is held at 40°C. The Calibration procedure is identical to that employed for the MFRSR using field measurements at the SGP site.

3. RESULTS

The instruments used in this study were located in the same general area of the ARM SGP central facility, but were spread over a distance of about 150 m. The AODs in Table 2 compare values for each instrument to the AATS-6, since it was expected to be the most accurate with a calibration just two weeks earlier at Mauna Loa. The nearest samples in time to the AATS-6 were used to derive bias and root-mean-squared differences. A first-order cloud filtering was achieved by comparing only with those AATS-6 AOD samples that yielded to an Angstrom parameter >0.75 . This resulted in as few as 461 samples to as many as 14,678 samples in the comparisons because of the different sampling strategies and days of operation. The AOD comparisons between AATS-6 and the other instruments are made for those filters whose central wavelengths match (for $\lambda = 380$ nm) or are within 6-8 nm ($\lambda = 519$ nm) of each other. For the other wavelengths the AODs were interpolated to the AATS-6 wavelength on a log vs. log AOD scale.

The AODs obtained from each instrument were derived independently of one another. Although the methods to remove Rayleigh, ozone and nitrogen dioxide optical depths may coincide in some instances, there was no attempt at a uniform reduction to aerosol optical depth from total optical depth. Consequently, some differences based on ozone assumptions or pressure corrections and the precise form of the Rayleigh and ozone calculations may add to the bias and rms differences shown in Table 2.

The best agreement with the AATS-6 overall was with the CIMEL leading to rms differences between 0.006 and 0.012. Cimel AODs tend to be slightly higher, 0.003 to 0.010, than the AATS-6 values (Table 2). Comparing AATS-6 and MFRSR leads to virtually no bias, but to slightly larger rms differences. Almost as good were results obtained between the AATS-6 and the RSS with the exception of the 380-nm channel of the RSS with a bias of 0.018 and rms difference of 0.025 optical depths.

Figures 1 and 2 are plots of the temporally averaged spectral aerosol optical depth on days with low aerosol loading (September 29) and moderate aerosol loading (October 2), respectively. The solid line in both figures is the RSS net optical depth after subtracting a Rayleigh scattering optical depth and an ozone optical depth. The broad peaks that remain are the result of absorption by water vapor, molecular oxygen, and collision-induced absorption by molecular oxygen pairs. As is clearly seen, filters in the other instruments are selected to avoid these bands in order to measure aerosol optical depths. The error bars on the AATS-6 data points include the uncertainty in Langley calibration, measured signal, airmass and gaseous absorption optical depths (Russell et al. 1993).

Figure 3, shows a time series of AOD obtained from the four radiometers on September 29 at wavelengths near 870 nm. The maximum AOD difference is 0.015. Figure 4 shows a time series of AOD of RSS and AATS-6 at the five AATS-6 aerosol wavelengths for September 27. Agreement throughout the day is 0.02 at 380 nm and 0.015 or better at the other wavelengths.

4. CONCLUSION

It should be pointed out that we have differenced the results in Table 2 with a single instrument (AATS-6) that was considered the most well-calibrated of the instruments, but which is not without error.

Consequently, some of the differences can be attributed to the reference instrument. The two instruments whose calibrations are tied to Langley calibrations at Mauna Loa and which rely on measuring the direct solar beam (CIMEL and AATS-6) agree most closely. The MFRSR and RSS, whose calibrations are based on the same robust technique employing 20 “on-site” Langley events, and which rely on subtracting the diffuse from the total component, agree closely with each other and agree within 0.025 with AATS-6.

In conclusion, we find that four well-calibrated instruments used in this comparison can retrieve AOD in the spectral range from 380 to 1020 nm with an accuracy of at least 0.025. Excluding the RSS results at 380 nm the accuracy is 0.015 close to the 0.01 AOD accuracy suggested as a goal by the WMO (1993). Encouraging is the fact that these accuracies were achieved for measurements at a site typical of where these measurements are required, rather than at an ideal mountain location.

Future attempts to repeat and improve this experiment should include calibrations immediately before or after the comparison. Reduction of total optical depths to aerosol optical depths should be based on identical algorithms applied to all of the data.

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TABLE 1. Central wavelengths () and bandwidths (, full widths at half maximum) of filtered instruments.

AATS-6		CIMEL		ARM MFRSR	
[nm]	[nm]	[nm]	[nm]	[nm]	[nm]
380.1	5.0	340	2	413.9	10
		380	4		
450.7	5.1	440	10		
525.3	5.0	500	10	499.3	10
				608.5	10
		670	10	665.1	10
863.9	5.3	870	10	859.9	10
941.4	5.8	940	10	938.0	10
1020.7	5.0	1020	10		

TABLE 2a: CIMEL minus AATS-6 (15 days, 461 pts)

Wavelength	Bias	RMS
380/380 nm	0.006 ± 0.011	0.012
451 nm	0.003 ± 0.008	0.008
525 nm	0.010 ± 0.006	0.012
864/870 nm	0.003 ± 0.009	0.009
1019/1020 nm	-0.001 ± 0.006	0.006

TABLE 2b: MFRSR minus AATS-6 (11 days, 14,678 pts)

Wavelength	Bias	RMS
451 nm	-0.002 ± 0.013	0.013
525 nm	0.000 ± 0.011	0.011
864/860 nm	0.002 ± 0.012	0.012

TABLE 2c: RSS minus AATS-6 (10 days, 4508 pts)

Wavelength	Bias	RMS
380 nm	-0.018 ± 0.018	0.025
451 nm	-0.009 ± 0.012	0.015
525 nm	-0.004 ± 0.009	0.010
864 nm	-0.006 ± 0.012	0.013
1020 nm	-0.004 ± 0.008	0.009

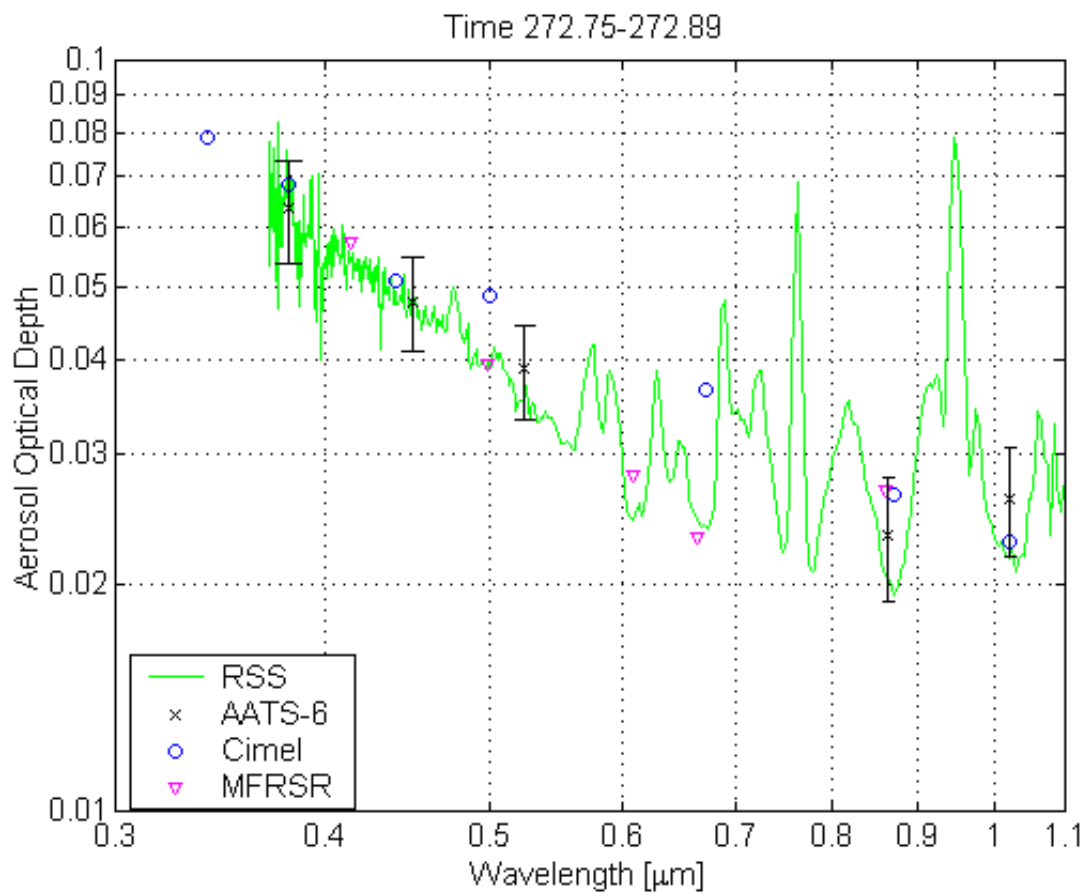


Fig. 1: Average aerosol optical depth (13:00-16:22 local time) versus wavelength on 29 September 1997; a low turbidity day.

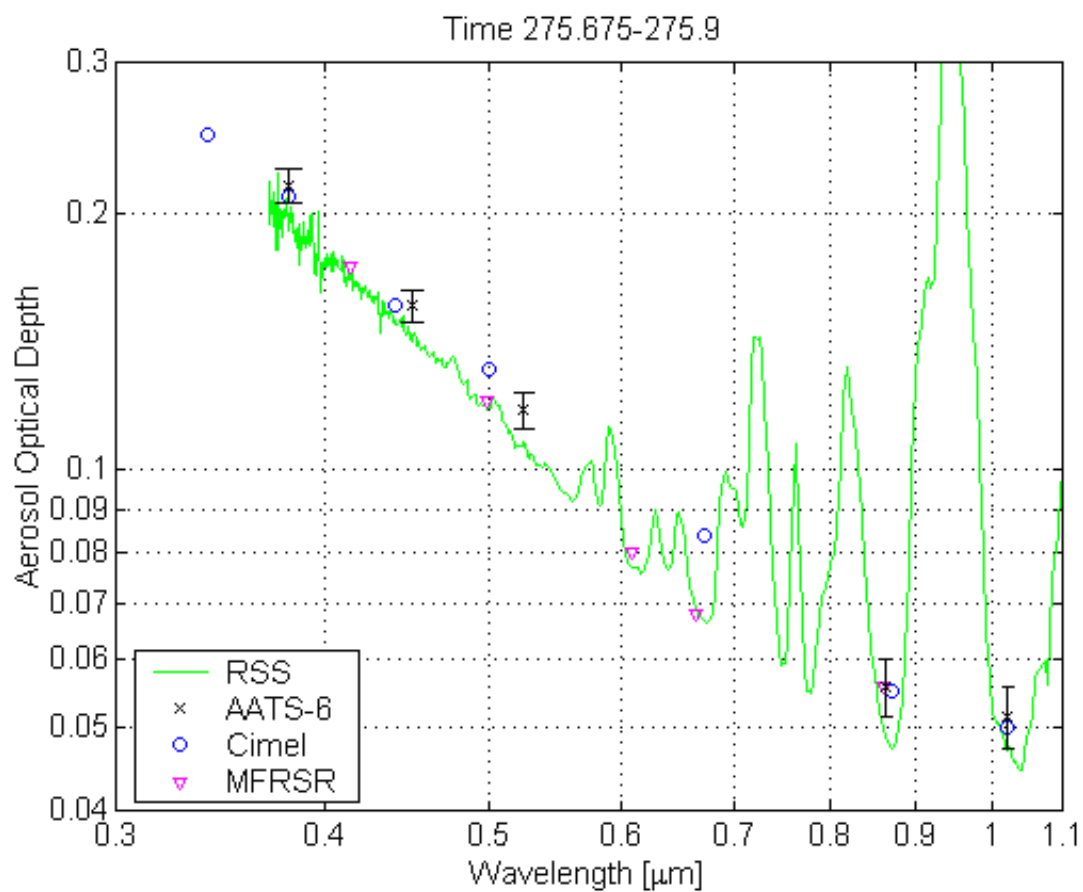


Fig. 2: Average aerosol optical depth (11:12-16:36 local time) versus wavelength on 2 October 1997; a moderately turbid day.

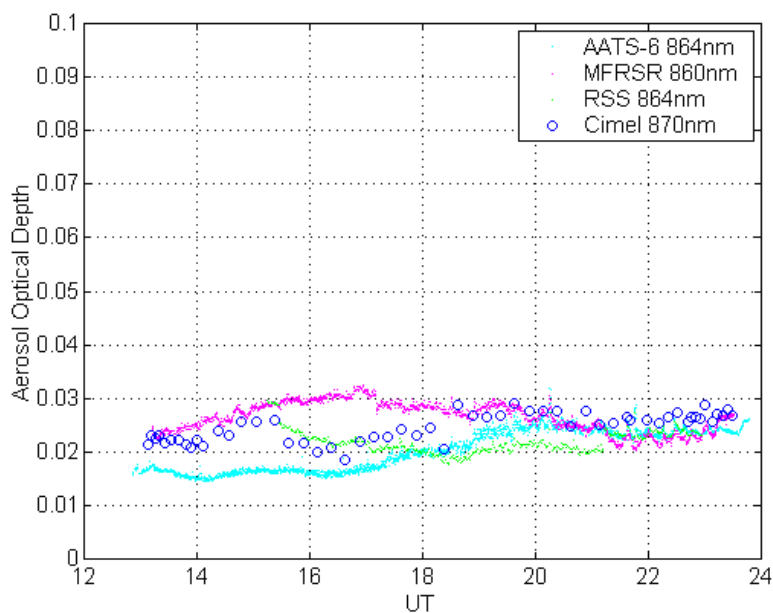


Fig. 3: Time series of aerosol optical depth near 870 nm as measured by 4 radiometers on September 29, 1997; a low turbidity day.

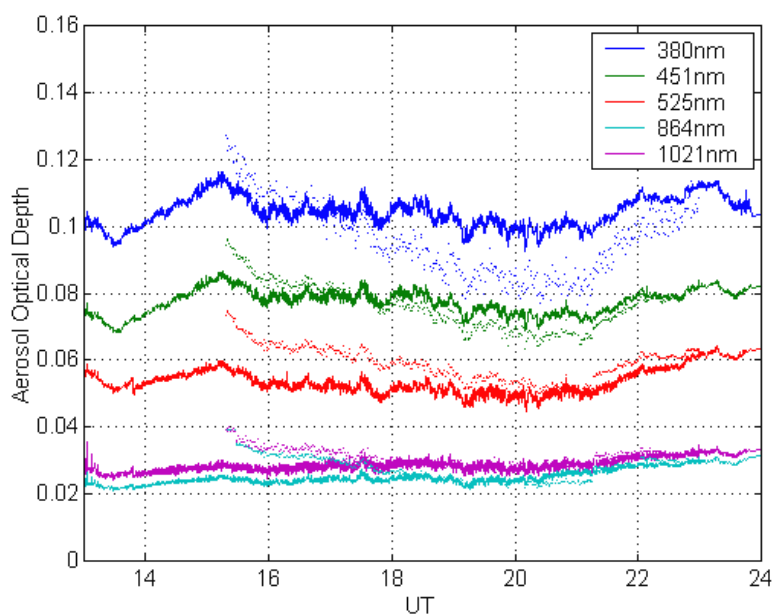


Fig. 4: Time series of aerosol optical depth of RSS (dots) and AATS-6 (lines) at the five AATS-6 aerosol wavelengths for September 27; a moderately turbid day.